

Unified Relativistic Gravitational Impedance Model: An Effective Dark Sector Fluid and its MOND Limit

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Abstract

We present a unified phenomenological model based on gravitational impedance $Z_{\text{eff}} = \eta G/c$. The central observation is that the relativistic mass flux $I_g = d(\gamma m_0)/dt = (m_0/c^2)\gamma^3(\vec{v} \cdot \vec{a})$ defines a specific potential $\Phi_g = Z_{\text{eff}} I_g$ whose sign is set by the scalar product $\vec{v} \cdot \vec{a}$. We show that (i) longitudinal acceleration $\vec{v} \cdot \vec{a} > 0$ yields $w \rightarrow -1$ (energy release, dark energy phase), (ii) longitudinal deceleration $\vec{v} \cdot \vec{a} < 0$ yields $w \rightarrow 0$ (energy absorption, dark matter phase), (iii) transverse motion gives no transfer at lowest order. The bare value $Z_g = G/c$ gives $\rho_g \sim 10^{-36} \text{ J/m}^3 \sim 10^{-27} \rho_\Lambda$ for the Milky Way, requiring amplification $\eta_\Lambda \sim 10^{27}$ for cosmic acceleration and $\eta_{a_0} \sim 10^6$ for the MOND acceleration $a_0 \simeq 1.2 \times 10^{-10} \text{ m/s}^2$. In Part II we derive $v_c^4 = GMa_0$ from the equilibrium $v_c^2 = \Phi_g$ with a_0 parametrized by Hubble boundary fluctuations $(\Delta V)^2/R_H$. The model does not replace ΛCDM but provides a single language for both dark components, with falsifiable predictions for $w(r)$ near jets and for Gaia streaming.

Contents

Part I – Relativistic Gravitational Impedance as an Effective Fluid	3
0.1 Introduction	3
0.2 Covariant Mass Flux	3
0.3 Impedance and Energy	3
0.4 Equation of State	3
0.5 Energetic Budget	3

Part II – MOND Limit and Tully-Fisher	4
0.6 Effective Mass Flux for a Galaxy	4
0.7 Deriving a_0	4
0.8 Tully-Fisher Relation	4
Part III – Tests, Limitations, Conclusion	4
0.9 Observational Tests	4
0.10 Limitations	4
0.11 Conclusion	5

Part I – Relativistic Gravitational Impedance as an Effective Fluid

0.1 Introduction

[Same references as before: Heaviside 1893, Sciama 1953, 1969, Unruh 1989, Milgrom 1983, Verlinde 2016]

The core idea, originally derived by the author from time differentiation of $m = \gamma m_0$, is

$$U \equiv Z_g \frac{dm}{dt}, \quad Z_g = \frac{G}{c}, \quad \frac{dm}{dt} \propto \vec{v} \cdot \vec{a} \quad (1)$$

where $\vec{v} \cdot \vec{a}$ is the scalar (not vector) product. For deceleration $U < 0$, suggesting absorption.

0.2 Covariant Mass Flux

Four-velocity u^μ , four-acceleration a^μ , projector $h_{\mu\nu} = g_{\mu\nu} + u_\mu u_\nu / c^2$,

$$I_g = \frac{m_0}{c^2} \gamma^3 (\vec{v} \cdot \vec{a}) = \frac{1}{c^3} g_{\mu\nu} u^\mu a^\nu m_0 \gamma^3 + O(\beta^4) \quad (2)$$

0.3 Impedance and Energy

$P = Z_{\text{eff}} I_g^2$, $Z_{\text{eff}} = \eta G / c$. For constant a :

$$E(\tau) = \frac{\eta G m_0^2 a^2}{8c^3} \left[\frac{\beta_\tau (\beta_\tau^2 + 1)}{(\beta_\tau^2 - 1)^2} + \frac{1}{2} \ln \left| \frac{\beta_\tau - 1}{\beta_\tau + 1} \right| \right] \quad (3)$$

Non-relativistic: $E \simeq \eta G m_0^2 a^2 \tau^3 / (3c^5)$.

0.4 Equation of State

Smooth interpolation avoiding manual sign:

$$w(\mathcal{F}) = -\frac{1}{2} [1 + \tanh(\mathcal{F}/\mathcal{F}_0)], \quad \mathcal{F} \equiv I_g \quad (4)$$

$w \rightarrow -1$ for $\mathcal{F} > 0$, $w \rightarrow 0$ for $\mathcal{F} < 0$.

$$T_{\text{imp}}^{\mu\nu} = \rho_g u^\mu u^\nu + w \rho_g c^2 h^{\mu\nu}, \quad \nabla_\mu (T_m^{\mu\nu} + T_{\text{imp}}^{\mu\nu}) = 0 \quad (5)$$

0.5 Energetic Budget

As computed in detail: Milky Way $P_{MW} \sim 2 \times 10^8 \eta$ W, $\rho_g \sim 4.3 \times 10^{-36} \eta$ J/m³, $\rho_g / \rho_\Lambda \sim 5 \times 10^{-27} \eta$. Requires $\eta_\Lambda \sim 2 \times 10^{26}$.

Part II – MOND Limit and Tully-Fisher

0.6 Effective Mass Flux for a Galaxy

Enclosed $M(r)$, dynamical time $t_{dyn} = r/v_c$, average longitudinal component $\langle \vec{v} \cdot \vec{a} \rangle = \epsilon v_c GM/r^2$,

$$\dot{M}_{\text{eff}} = \epsilon \Gamma \frac{M v_c}{c^2} \frac{GM}{r^2}, \quad \Phi_g = Z_g \dot{M}_{\text{eff}} = \epsilon \Gamma \frac{G^2 M^2 v_c}{c^3 r^2} \quad (6)$$

0.7 Deriving a_0

Define

$$a_0 \equiv \frac{c^3}{G^2 M} \Phi_g r = \epsilon \Gamma \frac{GM}{r^2} \frac{v_c}{c} \quad (7)$$

Parametrize $a_0 = \eta a_{\text{fluct}}$, $a_{\text{fluct}} = (\Delta V)^2/R_H$, $\Delta V = 494.6 \text{ km/s}$, $R_H = 123 \text{ Gly} = 1.16 \times 10^{27} \text{ m}$,

$$a_{\text{fluct}} = 2.11 \times 10^{-16} \text{ m/s}^2, \quad \eta_{a_0} = a_0/a_{\text{fluct}} \simeq 5.7 \times 10^5 \approx 10^6 \quad (8)$$

Using ΛCDM $R_H = 46.5 \text{ Gly}$ gives 5.3×10^{-16} and $\eta \sim 2.3 \times 10^5$, same order.

η interpreted as $\langle \epsilon \gamma^3 c/v_c \rangle$ integrated over halo, jets, hot gas. Single jet $\gamma \sim 10^3$ gives $\gamma^6 \sim 10^{18}$ in power.

0.8 Tully-Fisher Relation

From definition $\Phi_g = GMa_0/v_c^2$. Equilibrium $v_c^2 = \Phi_g$ gives

$$v_c^4 = GMa_0 \quad (9)$$

r cancels via definition of a_0 . Dimensionally $[GMa_0] = m^4/s^4 = [v^4]$.

Part III – Tests, Limitations, Conclusion

0.9 Observational Tests

- Lensing deficit $\Delta\Phi \sim Z_{\text{eff}} \dot{M}/c$ around jet hotspots - Euclid $\times$ LOFAR.
- Gaia DR3: correlation $v_c^4 - GMa_0 \propto \langle v_R v_\phi \rangle$.
- Lab: forced longitudinal acceleration with optical clocks, $P \sim 10^{-3} \text{ W}$ per kg with $\eta = 1$, $\sim 10^{24} \text{ W}$ with η_Λ - exclusion already.

0.10 Limitations

No action derivation, $w(\mathcal{F})$ ad-hoc, η scale-dependent, Pivot boundary non-standard.

0.11 Conclusion

Bare G/c insufficient by 10^{27} for dark energy, but with collective η provides unified language where acceleration/deceleration phases correspond to dark energy/dark matter. The MOND limit emerges as low-acceleration equilibrium.

Acknowledgments

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